



Recent Results in Rare Decays at BaBar

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Rare Decay Searches

Motivation: To search for rates enhanced over SM predictions that may arise from new virtual heavy mass particles in quantum loops

- 1. Search for $B \to K \nu \overline{\nu}$
- 2. Search for $B \to K^+ \tau^+ \tau^-$
- 3. Search for $B^0 \rightarrow \gamma \gamma$
- 4. Measurement of $B \rightarrow X_d \gamma$





The BaBar Dataset



BaBar collected 468 M BB pairs between 2000-2007 and 54 fb⁻¹ off-resonance data





The BaBar Detector



Search for $B \to K v \bar{v}$



Experiment BF (90% CL) Dataset Reference Belle < 1.4 x 10⁻⁵ 492 fb⁻¹ Chen etal PRL 99 221802, 2007 BaBar < 5.2 x 10⁻⁵ 82 fb⁻¹ Aubert et al. 94 1018011





Search for $B \to K_V \overline{v}$: Experimental Technique

Identify BB events by tagging with semi-leptonic $B \rightarrow D^{(*)} I_V$ decays







Search for $\,B \to K \nu \bar{\nu}\,$:Background Suppression



Nobs is from on-resonance data, Nbkg is expected from MC, ϵ is signal efficiency derived from MC

Mode	$\epsilon(\%)$	N_{sig}	N_{bkg}	N_{obs}	N_{excess}
K^+	0.16	2.9 ± 0.4	$17.6\pm2.6\pm0.9$	$19.4^{+4.4}_{-4.4}$	$1.8^{+6.2}_{-5.1}$
K_s^0	0.06	0.5 ± 0.1	$3.9\pm1.3\pm0.4$	$6.1^{+4.0}_{-2.2}$	$2.2^{+4.1}_{-2.8}$
$low-q^2$	0.24	2.9 ± 0.4	$17.6\pm2.6\pm0.9$	$19.4_{-4.4}^{+4.4}$	$1.8^{+\overline{6.2}}_{-5.1}$
high- q^2	0.28	2.1 ± 0.3	$187\pm10\pm46$	164_{-13}^{+13}	-23_{-48}^{+49}





Search for $B \rightarrow Kvv\overline{v}$:Results

arXiv:1009.1529



No signal observed but most stringent limits to date are set





Search for $B^+ \rightarrow K^+ \tau^+ \tau^-$



 $B^+ \rightarrow K^+ \tau^+ \tau \sim 50\%$ of total inclusive rate

Standard Model rate comparable to $\mu^+\mu^-$ or e^+e^- channels but new physics with a mass dependent coupling such as a Higgs in the Next-to-MSSM could enhance by $(m_{\tau}/m_{\mu})^2 \sim 280$ (G.Hiller PRD 70 034018 (2004))





Search for $B^+ \rightarrow K^+ \tau^+ \tau^-$

Analysis Technique



Search for $B^+ \rightarrow K^+ \tau^+ \tau^-$: Backgrounds







Search for $B^+ \rightarrow K^+ \tau^+ \tau^-$: Results



Systematic	%
B Counting	1.1
Tag Efficiency	3.2
Signal Efficiency	14.8
Background Estimation	17.3

Expected Bkgd: 64.7 +/- 7.3

Data Events: 47

 $B (B^+ \rightarrow K^+ \tau^+ \tau^-) < 0.0033 (90\% CL)$

(First limit to date)





Search for $B^0 \rightarrow \gamma \gamma$

Standard Model \bar{B}^0 u, c, t $W^ B(B^0 \rightarrow \gamma \gamma) \sim 3 \times 10^{-8}$
(Bosch and Buchalla,, JHEP 0208:054 (2002)) \bar{d} $W^ B(B^0 \rightarrow \gamma \gamma) \sim 0208:054 (2002)$ \bar{d} Physics Beyond Standard Modelb $W^ B(B^0 \rightarrow \gamma \gamma) \sim O(10^{-7})$
Aliev and Turin, PRD 58 095014
(2HDM models or R-parity violating SUSY)

Experimental constraints from $b \rightarrow d\gamma$ experiment

Previous Measurements

Experiment	BF (90% CL)	Dataset	Reference	
L3	< 1.9 x 10 ⁻⁵	2.95x10 ⁶ (Z→had)	Acciarri et al. Phys. Lett. B, 363, 1995	
BaBar	< 1.7 x 10 ⁻⁶	19 fb ⁻¹	Aubert et al. PRL 87, 24, 2001	
Belle	< 6.1 x 10 ⁻⁷	104 fb ⁻¹	Villa et al. PRD 73, 2006	
ABAR.	Colin Jessop at Heavy Quarks and Leptons			



$B^0 \rightarrow \gamma \gamma$ Backgrounds





$$B(B \rightarrow \gamma \gamma) = (1.7 \pm 1.1(stat.) \pm 0.2(sys.)) \times 10^{-7} \quad (1.9\sigma significance)$$







$$B(B \rightarrow \gamma \gamma) < 3.3 \times 10^{-7} at 90\% C.L.$$



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arXiv:1010.2229

Measurement of $B \rightarrow X_d \gamma$ and Extraction of $|V_{td}/V_{ts}|$





Measurement with penguins to search for New Physics

Previously used ratio of exclusives ($\rho,\omega\gamma/K^*\gamma)$ but limited by form factor uncertainty

Inclusive method is theoretically cleaner

Use the sum-of-exclusives technique (~50% of modes covered. Largest systematic from missing modes)

$$B \rightarrow X_{d}\gamma \qquad B \rightarrow X_{s}\gamma$$

$$B^{0} \rightarrow \pi^{+}\pi^{-}\gamma \qquad B^{0} \rightarrow K^{+}\pi^{-}\gamma$$

$$B^{+} \rightarrow \pi^{+}\pi^{0}\gamma \qquad B^{+} \rightarrow K^{+}\pi^{0}\gamma$$

$$B^{+} \rightarrow \pi^{+}\pi^{-}\pi^{+}\gamma \qquad B^{+} \rightarrow K^{+}\pi^{-}\pi^{+}\gamma$$

$$B^{0} \rightarrow \pi^{+}\pi^{-}\pi^{0}\gamma \qquad B^{0} \rightarrow K^{+}\pi^{-}\pi^{0}\gamma$$

$$B^{0} \rightarrow \pi^{+}\pi^{-}\pi^{+}\pi^{-}\gamma \qquad B^{0} \rightarrow K^{+}\pi^{-}\pi^{+}\pi^{-}\gamma$$

$$B^{+} \rightarrow \pi^{+}\pi^{-}\pi^{+}\pi^{0}\gamma \qquad B^{+} \rightarrow K^{+}\pi^{-}\pi^{+}\pi^{0}\gamma$$

$$B^{+} \rightarrow \pi^{+}\eta\gamma \qquad B^{+} \rightarrow K^{+}\eta\gamma$$



Measurement of $B \rightarrow X_d \gamma$ and Extraction of $|V_{td}/V_{ts}|$

C

 $471M B\overline{B}$





Measure for $M_{X_{ds}} < 2.0 \ GeV^2$ $\frac{\Gamma(B \to X_d \gamma)}{\Gamma(B \to X_s \gamma)} = 0.040 \pm 0.009(stat.) \pm 0.010(sys.)$ Correct for unmeasured $M_{X_{dx}} > 2.0 \ GeV^2$ using Kagan & Neubert (*PRD* 58 094012) spectrum with $m_{\rm h} = 4.65 \pm 0.05 \ \mu_{\pi}^2 = -0.52 \pm 0.08 \ (HFAG)$ Extract $\left|\frac{V_{td}}{V}\right|$ using the calculations of Ali,Asatrian & Greub using β as input rather than (ρ, η) (Phy. Lett. B 429 87 (1998)) $= 0.199 \pm 0.022(\text{stat.}) \pm 0.024(\text{sys.}) \pm 0.002(\text{th.})$ $\frac{V_{td}}{V_{t}}$ Belle $(\rho, \omega)\gamma$ $0.195^{+0.020}_{-0.019}\pm0.015$ BaBar $(\rho, \omega)\gamma$ 0.233+0.025+0.022 Average $(\rho, \omega)\gamma$ $0.210 \pm 0.015 \pm 0.018$ arXiv 1005.4087v1 PRD 82:051101 2010 BaBar X, y 0.199 ± 0.032 ± 0.001 **B Mixing Average** Radiative Decay Avg 0.206 ± 0.019 UNIVERSITY OF NOTRE DAME 0.45 0.05 0.1 0.15 0.2 0.25 0.3 0.35 0.4 $|V_{td}/V_{ts}|$

BaBar continues to mine its dataset for evidence of physics beyond the Standard Model in rare B decays:

Most stringent limits presented for

 $B(B \rightarrow K \nu \nu) \le 1.4 \text{ x } 10^{-5}$

 $B(B \rightarrow K^+ \tau^+ \tau^-) < 3.3 \text{ x } 10^{-3}$

 ${\rm B}({\rm B}^{0} \rightarrow \gamma\gamma ~) < 3.3 ~ x ~ 10^{\text{--}7}$

Measurement of $\frac{\Gamma(B \to X_d \gamma)}{\Gamma(B \to X_s \gamma)} = 0.040 \pm 0.009(stat.) \pm 0.010(sys.)$ $\frac{\left|\frac{V_{td}}{V_{ts}}\right| = 0.199 \pm 0.022(stat.) \pm 0.024(sys.) \pm 0.002(th.)$

Techniques developed will hopefully be used to observe these modes at SuperB !



