

## Softness of doubly-magic $^{78}\text{Ni}$ and related topics <sup>4</sup>

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Neutron-rich nickel isotopes in the vicinity of  $^{78}\text{Ni}_{50}$  are currently in the focus of modern nuclear physics and astrophysics studies. The interest in this region is motivated by the doubly magic nature of  $^{78}\text{Ni}$  and understanding the way in which the neutron excess will affect the properties of nearby nuclei and nucleon-nucleon interaction. The astrophysical importance is related to the understanding of the nuclear mechanism of the rapid capture of neutrons by seed nuclei through the r-process. The path of this reaction network is expected in neutron-rich nuclei for which there is little experimental data, and the precise trajectory is dictated by the details of the shell structure far from stability.

The shell-model orbitals for neutrons in nuclei with  $Z = 28$  and  $N=28-50$  ( $^{56}\text{Ni}$ - $^{78}\text{Ni}$ ) are the same as those for protons in nuclei with  $N=50$  and  $Z=28-50$  ( $^{78}\text{Ni}$ - $^{100}\text{Sn}$ ). However the comparison of the properties of the  $N=50$  isotones near  $^{100}\text{Sn}$  and Ni-isotopes near  $^{78}\text{Ni}$  with a structure dominated by the  $g_{9/2}$ -orbital configurations indicate a considerable difference in the corresponding spectra. New effective shell model interactions [1] links the difference to the unusual enhancement of the two-body interaction in  $J^\pi = 2^+$  and  $J^\pi = 4^+$  channels for neutron  $g_{9/2}$  orbital near  $^{78}\text{Ni}$  [2]. This is well illustrated by experimental data for essentially two-nucleon hole states in  $^{76}\text{Ni}$  and  $^{98}\text{Cd}$  nuclei.

In present contribution we discuss the possibility to understand the observed differences and unusual properties of the interaction by breaking the  $^{56}\text{Ni}$  core. An issue of  $^{56}\text{Ni}$  core breaking effects for the structure and beta decay of neutron rich nickel isotopes will be addressed.

[1] A.F. Lisetskiy, B. A. Brown, M. Horoi, H. Grawe, *Phys. Rev. C* **70**,044314,(2004).

[2] M. Horoi, A. F. Lisetskiy, B. A. Brown, AIP. Conf. Proc. **764**, 170 (2005).

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